Sources and Deposition of Reactive Gaseous Mercury in the Marine Atmosphere

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Sources and deposition of reactive gaseous mercury in the marine atmosphere

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Abstract

Observations of reactive gaseous mercury (RGM) in marine air show a consistent diurnal cycle with minimum at night, rapid increase at sunrise, maximum at midday, and rapid decline in afternoon. We use a box model for the marine boundary layer (MBL) to interpret these observations in terms of RGM sources and sinks. The morning rise and midday maximum are consistent with oxidation of elemental mercury (Hg 0) by Br atoms, requiring <2 ppt BrO in most conditions. Oxidation of Hg 0 by Br accounts for 35–60% of the RGM source in our model MBL, with most of the remainder contributed by oxidation of Hg 0 by ozone (5–20%) and entrainment of RGM-rich air from the free troposphere (25–40%). Oxidation of Hg 0 by Cl is minor (3–7%), and oxidation by OH cannot reproduce the observed RGM diurnal cycle, suggesting that it is unimportant. Fitting the RGM observations could be achieved in the model without oxidation of Hg 0 by ozone (leaving Br as the only significant oxidant) by increasing the entrainment flux from the free troposphere. The large relative diurnal amplitude of RGM concentrations implies rapid loss with a lifetime of only a few hours. We show that this can be quantitatively explained by rapid, mass-transfer-limited uptake of RGM into sea-salt aerosols as HgCl 3 and HgCl 4. Our results suggest that 80–95% of Hg 0 in the MBL should be present in sea-salt aerosol rather than gas-phase, and that deposition of sea-salt aerosols is the major pathway delivering Hg 0 to the ocean.

1. Introduction

Atmospheric deposition is the principal source of mercury to the ocean (Lindberg et al., 2007). Human industry has increased this source by about a factor of three relative to natural conditions (Mason and Sheu, 2002), resulting in a corresponding increase in the surface ocean pool and raising concerns about mercury accumulation in marine biota (Mergler et al., 2007). Mercury is emitted to the atmosphere mainly as elemental mercury (Hg 0) and is oxidized within the atmosphere to Hg II, which is highly water-soluble and the major depositing form (Lindberg et al., 2007). A critical step towards understanding the deposition of mercury to the ocean is quantifying the supply of Hg II to the atmospheric marine boundary layer (MBL), which extends about 1 km above the ocean surface and is in turbulent contact with it. We apply here a box model for the MBL to analyze several recent data sets of reactive gaseous mercury (RGM, representing gaseous Hg II) over the North Atlantic and Pacific Oceans (Laurier et al., 2003; Jaffe et al., 2005; Laurier and Mason, 2007), in order to better understand the Hg II budget in the MBL and the implications for ocean uptake.

Mercury has an atmospheric lifetime of 0.5–1 years, enabling transport on a global scale (Lindberg et al., 2007). Oxidation of Hg 0 must be photochemical, as evidenced by the observed seasonal cycle of Hg 0 (Selin et al., 2007) and by the diurnal cycle of RGM (Laurier and Mason, 2007). But it is still unclear which oxidants dominate. Standard models assume gas-phase OH and ozone to be the main oxidants (Shia et al., 1999; Petersen et al., 2001; Cohen et al., 2004; Lin et al., 2006; Seigneur et al., 2006; Selin et al., 2007; Sillman et al., 2007). However, recent thermodynamic and kinetic analyses suggest that these reactions are extremely slow in the atmosphere (Calvert and Lindberg, 2005; Hynes et al., 2008). Gas-phase oxidation by Br atoms is known to cause rapid loss of Hg 0 in the Arctic boundary layer in spring, an environment with particularly active bromine chemistry (Goodsite et al., 2004; Brooks et al., 2006; Tackett et al., 2007). Holmes et al. (2006) suggested that Br atoms in the free troposphere (above the boundary layer) could provide the main global sink of Hg 0.

Mercury–bromine chemistry could also be active within the MBL at low and middle latitudes. Mason and Sheu (2002) proposed that 40% of Hg 0 emitted from the ocean is oxidized within the MBL and deposited back; they implicated halogen oxidants but did not...
offer a specific mechanism. A direct test of Hg–halogen chemistry at a midlatitude coastal site by simultaneous BrO and RGM measurements was inconclusive due to the high BrO detection limit (2 pptv) (Keene et al., 2007; Laurier and Mason, 2007), but Br could still be important as a Hg0 oxidant at lower BrO concentrations, as demonstrated below. Hedgcock et al. (2003, 2004, 2005, 2006) showed that a box model for the MBL with atomic Br released from sea-salt aerosols could reproduce the mean concentrations and diurnal amplitudes of RGM observed over the Mediterranean Sea. However, the Hg + Br reaction kinetics that they used (Ariya et al., 2002) seem too fast in light of recent data (Goodsite et al., 2004; Balabanov et al., 2005; Donohoue et al., 2006; Ariya et al., 2008). As a result, their model required a diurnal cycle in marine Hg0 emission to balance oxidation and thus account for the observed lack of diurnal variation in atmospheric Hg0 concentrations (Hedgcock and Pirrone, 2004). A diurnal cycle of Hg0 emission is difficult to justify, even with photoreduction of HgII in surface water, because the gaseous ventilation time of the marine photic zone is ~1 month (Strode et al., 2007). We will show that these shortcomings can be corrected while preserving Hedgcock’s central result that Br is a major Hg0 oxidant in the MBL.

2. Observed RGM diurnal variability

We analyze reactive gaseous mercury (RGM) data from cruises in the remote Atlantic (Laurier and Mason, 2007) and Pacific Oceans (Laurier et al., 2003), and from coastal measurements in Okinawa, Japan (Jaffe et al., 2005). RGM is the measured gas-phase component of HgII, collected here using a KCl denuder (Landis et al., 2002). The Pacific data were collected between Osaka, Japan and Honolulu, Hawaii over 20 days during May–June 2002. We separated the midlatitude and subtropical data at 27°N based on a sharp rise in water and air temperatures. The Atlantic data were collected in subtropical waters between Bermuda and Barbados over 11 days during August–September 2003. In both data sets the RGM concentrations are 2-h averages and the detection limit for individual measurements is 3 pg m⁻³. We removed data contaminated by the ship’s exhaust, which occurred on only one day during the Pacific cruise. The ships also measured Hg0, O₃, and CO concentrations in addition to standard oceanographic and meteorological variables. Laurier et al. (2003) and Laurier and Mason (2007) provide further details on the cruise tracks and sampling methods.

The Okinawa data span 40 days during March–May 2004. They were collected at the northern tip of the island (26.8° N, 128.2° E, 60 m a.s.l.) with prevailing onshore flow. The site has minimal impact from anthropogenic sources on the island. Averaging times for the measurements were 5 min for Hg0 and 3 h for RGM and particulate mercury (radii < 1.3 µm). The RGM detection limit for individual measurements is 1 pg m⁻³. Hg0 at Okinawa is strongly correlated with CO, indicating an East Asian industrial origin, but RGM is not (Jaffe et al., 2005; Strode et al., 2008). We discard 5 days of observations with the greatest pollution transport from continental Asia.

The dominant feature of RGM temporal variability is its diurnal cycle, as shown in Fig. 1. The diurnal cycle is statistically significant (p = 0.05 from Kruskal–Wallis test of constant mean) at Okinawa and in the subtropical Atlantic and Pacific data, but is much weaker and does not reach statistical significance in the midlatitude Pacific data (p = 0.2). RGM increases rapidly at sunrise, peaks at midday, declines through the afternoon, and falls near detection limit at night. Neither deposition nor boundary layer growth can drive this diurnal cycle, since the winds and MBL depth have little diurnal variability (Hignett, 1991). Rather, RGM production due to changing oxidant levels must be responsible. Although the amplitude of the RGM diurnal cycle varies between the data sets in Fig. 1, the consistent shape and phase argue for a common driving mechanism.

The RGM increase at sunrise suggests a photochemical oxidant produced by photolysis of weakly-bound nighttime reservoirs. Br and Cl can exhibit this behavior (von Glasow et al., 2002). By contrast, OH in the MBL increases only slowly after sunrise and peaks at midday (Nowak et al., 2001). Selin et al. (2007) previously noted that Hg0 oxidation by OH yields an afternoon peak of RGM delayed by several hours relative to the observations at Okinawa. In contrast to RGM, the Hg0 concentrations in all data sets show no significant diurnal cycle, which places an upper bound on the Hg0 oxidation rate. Given 10% measurement uncertainty for Hg0 (Jaffe et al., 2005), daytime oxidation must be less than 3–5% d⁻¹ or 4–8 pg m⁻³ h⁻¹.

The relative diurnal amplitude of RGM, expressed as the difference between maximum and minimum concentrations divided by the 24-h mean, ranges from 1 to 3 depending on the data set. Concentrations decrease by 40–70% over the 12–18 Local Solar Time (LST) period, implying a lifetime of a few hours at most. Nighttime RGM observations tend to decrease as wind speed increases (Fig. 2), suggesting a fast deposition sink rather than chemical loss. Hedgcock and Pirrone (2001) and Selin et al. (2007) previously speculated that uptake by sea-salt aerosols might provide such a fast deposition sink, and we provide further mechanistic support for this hypothesis below.

3. Model description

We interpret the observed diurnal cycles of RGM with a box model of the MBL applied to the different data sets of Fig. 1. The model calculates the concentration of gaseous HgII (c, mol m⁻³) based on chemical production and loss (P and L, mol m⁻³ s⁻¹), dry deposition to the ocean surface (Fl, mol m⁻² s⁻¹), entrainment at the top of the MBL (Fe, mol m⁻³ s⁻¹), and uptake by sea-salt aerosols (J, mol m⁻³ s⁻¹):

\[ dc = \frac{P - L + \frac{P_{e} - F_{e}}{Z} - J}{\tau} \]

where Z is the MBL depth. Downward fluxes are positive. For comparison with observations we assume that all gas-phase HgII species are measured as RGM.

Table 1 summarizes the model parameters, which are taken from measurements when available. Other parameters are typical MBL values, which suffice to compare proposed alternative HgII formation and loss pathways against RGM data. We choose the concentrations of Br and free-tropospheric RGM to best match the amplitude and mean of the RGM diurnal cycle in each data set, as described below, then evaluate these concentrations against the few available measurements.

Table 2 lists the kinetic data used to calculate P and L. Oxidation to Hg0 by halogen atoms and OH proceeds in two addition reactions through the intermediate HgX (X = Br, Cl, OH). We neglect oxidation initiated by OH in our reference model due to the instability of HgOH under atmospheric conditions (Goodsite et al., 2004; Calvert and Lindberg, 2005; Hynes et al., 2008), but we include it in a sensitivity test.

The entrainment flux at the top of the boundary layer is the product of an assumed entrainment velocity \( v_{e} = 0.5 \text{ cm s}^{-1} \) (Faloona et al., 2005) and the concentration difference between the MBL and the free troposphere (CFR):

\[ F_{e} = v_{e}(C_{FR} - c). \]

The dry deposition flux is calculated by applying a deposition velocity (\( v_{d} \)) to the RGM concentration:
Due to the high solubility of Hg\textsuperscript{II} species in water (Table 2), \(v_d\) is limited by aerodynamic resistance. For neutrally stable conditions typical of the marine atmosphere, \(v_d(z) = \frac{ku\ast}{\ln(z/z_0)}\) (4), where \(k = 0.4\) is the von Kármán constant (Seinfeld and Pandis, 2006). Most aerodynamic resistance to deposition occurs in the lowest few meters, so we use a reference height \(z = 10\) m at which wind speeds were measured. The roughness length \((z_0)\) and friction velocity \((u\ast)\) are calculated from the observed wind speed at 10 m \((u(z))\) by iteratively solving the system of equations (Stull, 1988, pp. 377, 381):

\[
z_0 = \frac{\alpha_C u^2}{g}
\]

(5)

\[
u\ast = \frac{ku(z)}{\ln(z/z_0)}
\]

(6)

where \(\alpha_C = 0.016\) is the Charnock constant and \(g = 9.8\) m s\(^{-2}\) is the gravitational acceleration.

We find \(v_d = 0.4\)–1.4 cm s\(^{-1}\) for the different data sets (Table 1). Combined with a typical MBL depth \(Z = 750\) m, these deposition rates imply an RGM lifetime of 15–50 h against dry deposition. This is too slow to explain the rapid decrease observed in the afternoon and evening hours (Fig. 1). A more important pathway for RGM loss may be uptake by sea-salt aerosols followed by aerosol deposition. We assume that all Hg\textsuperscript{II} species have the same solubility as HgCl\(_2\). Once dissolved in aqueous aerosols or cloud droplets, Hg\textsuperscript{II} species convert to HgCl\(_2\) due to the abundance of Cl\(^-\) (Hedgecock and Pirrone, 2001). Complexation of HgCl\(_2\) to HgCl\(_3^+\) and HgCl\(_4^{2-}\) preferentially fractionates HgCl\(_2\) into the sea-salt aerosol (Table 2) (Hedgecock and Pirrone, 2001). The equilibrium partitioning depends on the sea-salt aerosol chloride concentration, which is a function of the water vapor saturation ratio \((S)\):

\[
\text{[Cl}^-\text{]} = \frac{\rho_{SS} f_C}{M_{Cl}} \left(\frac{1 - S}{2 - S}\right)^{3.7} \left(\frac{4}{3}\right)^3,
\]

(7)

where \(M_{Cl} = 35.5\) g mol\(^{-1}\) is the atomic weight of Cl, \(\rho_{SS} = 2200\) kg m\(^{-3}\) is the density of dry sea salt, and \(f_C = 0.55\) is the mass fraction of Cl in dry sea salt. This formula follows from an empirical expression for sea-salt aerosol growth relative to its dry radius \((r_{dry})\) (Lewis and Schwartz, 2006):

\[
r = r_{dry} \left[\frac{2 - S}{1 - S}\right]^{1/3}.
\]

(8)

The relation holds over \(0.45 < S < 0.99\) which spans the range found in the MBL. For typical marine conditions of \(S = 0.8\) and sea-salt aerosol liquid water content \(L = 1 \times 10^{-10}\) m\(^3\) water m\(^{-3}\) air, it follows that \([\text{Cl}^-]\) = 4.5 M and 85% of total Hg\textsuperscript{II} is dissolved in sea-salt aerosol at equilibrium.
Because of the short sea-salt aerosol lifetime, one cannot assume RGM to be in equilibrium with the aerosol; RGM uptake is in fact limited by mass transfer. We find that this limitation slows down RGM loss by a factor of 1.5–3 relative to assuming equilibrium. Let \( c_0 \) (mol m\(^{-3}\)) denote the aqueous-phase concentration of Hg\(^{2+}\) = HgCl\(_2\)(aq) + HgCl\(_3\) + HgCl\(_4\)\(^2-\), and let \( H \) be the dimensionless effective Henry’s law constant for HgCl\(_3\) accounting for the polychloride complexes. The net uptake of Hg\(^{2+}\) by the aerosol is given by

\[
J = L k_{\text{mt}} \left( c - c_0 \right) \frac{a}{H},
\]

where \( k_{\text{mt}} \) is a mass-transfer coefficient to the aerosol. For a monodisperse aerosol with number concentration \( N \) (m\(^{-3}\)) of particles of radius \( r \), \( k_{\text{mt}} \) can be approximated by (Sander, 1999)

\[
k_{\text{mt}} = \frac{4 \pi r^2}{L} \left( \frac{r}{D_k} + \frac{4}{r v_a} \right)^{-1} N,
\]

where \( D_k = 0.1 \text{ cm}^2 \text{ s}^{-1} \) is the diffusivity of RGM in air, \( v_a = 1.5 \times 10^4 \text{ cm s}^{-1} \) is the mean molecular speed, and \( a = 0.5 \) is the assumed accommodation coefficient for HgCl\(_2\). Results are insensitive to \( a \) in the range 0.1–1 typical of highly soluble gases.

We calculate RGM uptake by sea-salt aerosol by considering \( m = 5 \text{ dry aerosol size classes with radius divisions at 1, 3.2, 5.4, 7.6, 9.8, and 12 \mu m} \). Through sensitivity tests we found that smaller, submicron aerosols are insignificant sinks of RGM because of their small liquid water content while aerosols larger than 12 \( \mu m \) are also insignificant sinks because of their small number and short atmospheric residence time. For each size class, we calculate the number concentration by integrating over the steady-state aerosol number distribution governed by the balance between aerosol production (Lewis and Schwartz, 2004, p. 321 lognormal approximation) and loss by dry deposition (Zhang et al., 2001). For simulations over the Mediterranean, Hedgecock et al. (2005) assumed fixed \( L = 3 \times 10^{-11} \text{ m}^2 \text{ s}^{-1} \text{ m}^3 \text{ air} \) and an aerosol residence time of 3 days. Using the wind speed from that study (3.8 m s\(^{-1}\)) and \( S = 0.8 \), we obtain similar values of \( L = 4.7 \times 10^{-11} \text{ m}^2 \text{ s}^{-1} \text{ m}^3 \text{ air} \) and a residence time of 1.8 days averaged over all size classes. The total aerosol uptake in Equation (1) is a sum over the fluxes for each size class:

\[
J = \sum_j m_{\text{air}} J_j.
\]

In addition to Equation (1), the model calculates the sea-salt aerosol Hg\(^{2+}\) (aq) concentration:

\[
\frac{\text{d}c_{\text{aq},i}}{\text{d}t} = f_i \left( L_i - v_{d,i} c_{\text{aq},i} \right).
\]

Here \( c_{\text{aq},i} \), \( L_i \), and \( v_{d,i} \) are the Hg\(^{2+}\) (aq) concentration, liquid water content, and aerosol deposition velocity for the \( i \)th size class. This

---

**Table 1**

<table>
<thead>
<tr>
<th>Parameter</th>
<th>Region</th>
<th>Okinawa</th>
<th>Pacific cruise midlatitudes</th>
<th>Pacific cruise subtropics</th>
<th>Atlantic cruise subtropics</th>
</tr>
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<tr>
<td>[Hg(^{0})] [^{\text{a}}] (ng m(^{-3}))</td>
<td>2.0</td>
<td>2.5</td>
<td>2.5</td>
<td>1.6</td>
<td>13</td>
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<tr>
<td>[Cl(^{-})], ppb</td>
<td>31</td>
<td>35</td>
<td>10</td>
<td>13</td>
<td>13</td>
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<td>Temperature, K</td>
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<td>283</td>
<td>298</td>
<td>301</td>
<td>301</td>
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<td>10-m wind speed, m s(^{-1})</td>
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<td>5.9</td>
<td>5.1</td>
<td>11.2</td>
<td>75</td>
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<tr>
<td>Relative humidity, %</td>
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<td>77</td>
<td>84</td>
<td>75</td>
<td>75</td>
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<tr>
<td>Derived [^{\text{b}}]</td>
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<tr>
<td>( v_a ), cm s(^{-1})</td>
<td>0.38</td>
<td>0.56</td>
<td>0.46</td>
<td>1.4</td>
<td>1.4</td>
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<tr>
<td>([\text{Cl}^{-}]), M</td>
<td>4.5</td>
<td>5.1</td>
<td>3.8</td>
<td>5.4</td>
<td>5.4</td>
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<tr>
<td>( H ), M atm(^{-1})</td>
<td>2.6 \times 10(^7)</td>
<td>3.2 \times 10(^7)</td>
<td>1.8 \times 10(^8)</td>
<td>3.7 \times 10(^8)</td>
<td>3.7 \times 10(^8)</td>
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<tr>
<td>( L ), m(^{3}) \text{ air} \text{ m}^2 \text{ s}^{-1} \</td>
<td>8.3 \times 10(^{-10})</td>
<td>1.6 \times 10(^{-10})</td>
<td>1.3 \times 10(^{-10})</td>
<td>5.3 \times 10(^{-10})</td>
<td>5.3 \times 10(^{-10})</td>
</tr>
<tr>
<td>( F_v ), m(^{3}) \text{ air} \text{ m} \text{ s}^{-1} \</td>
<td>3.6 \times 10(^{-13})</td>
<td>1.1 \times 10(^{-12})</td>
<td>7.1 \times 10(^{-13})</td>
<td>8.8 \times 10(^{-12})</td>
<td>8.8 \times 10(^{-12})</td>
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<tr>
<td>Hg(^{0}) lifetime, d</td>
<td>120</td>
<td>140</td>
<td>90</td>
<td>10</td>
<td>10</td>
</tr>
<tr>
<td>RGM, lifetime, h</td>
<td>8.1</td>
<td>3.8</td>
<td>6.7</td>
<td>0.7</td>
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<tr>
<td>Other [^{\text{c}}]</td>
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<td>( Z ), m</td>
<td>750</td>
<td>750</td>
<td>750</td>
<td>750</td>
<td>750</td>
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<tr>
<td>([\text{Br}]), cm(^{-3})</td>
<td>4.3 \times 10(^7)</td>
<td>2.0 \times 10(^7)</td>
<td>8.7 \times 10(^6)</td>
<td>4.3 \times 10(^6)</td>
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<td>([\text{Cl}]), cm(^{-3})</td>
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<td>1.5 \times 10(^8)</td>
<td>1.5 \times 10(^8)</td>
<td>1.5 \times 10(^8)</td>
<td>1.5 \times 10(^8)</td>
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<tr>
<td>([\text{OH}]), cm(^{-3})</td>
<td>1.1 \times 10(^6)</td>
<td>6.3 \times 10(^5)</td>
<td>1.2 \times 10(^6)</td>
<td>1.1 \times 10(^6)</td>
<td>1.1 \times 10(^6)</td>
</tr>
<tr>
<td>( \text{GCPG, pg m}^{-3} )</td>
<td>10</td>
<td>20</td>
<td>20</td>
<td>20</td>
<td>20</td>
</tr>
</tbody>
</table>

\(^{\text{a}}\) Means over the observation period.

\(^{\text{b}}\) We display \( H \) in conventional units here, so one should multiply this by 24 atm M\(^{-1}\) to obtain the dimensionless form used in Equation (9). \( F_v \) is the volumetric production flux of sea-salt aerosols from the ocean surface. Lifetimes are calculated with the 24-h mean fluxes in Fig. 4. The RGM lifetime accounts for ventilation, while the Hg\(^{0}\) lifetime includes only MBL chemical loss. Other derived parameters are calculated from the measurements with the equations in Section 3.

\(^{\text{c}}\) Other parameters are from literature and models as described in Section 3. [Br\(^{-}\)], [Cl\(^{-}\)], and [OH\(^{-}\)] are 24-h means which multiply the relative amplitudes in Fig. 3.

---

**Fig. 2.** Nighttime RGM concentrations (0–5 LST) vs. wind speed. Symbols show observed means for individual days. Lines show the relationship predicted by the box model for each data set by varying wind speed while holding other parameters constant.
The concentration of HgII in our model is the sum of RGM and particulate HgII. We use a dissociation rate coefficient of HgCl₂, $k_1$, to form HgCl⁻, which is the ideal gas constant and $k_1$ is the rate coefficient for HgCl⁻ + Cl⁻ → HgCl₂ (from the table above). We find a dissociation rate coefficient of HgCl₂, $k_1$, in units of M atm⁻¹ s⁻¹ for HgCl₂, which is negligibly slow.

We assume the same solution for all HgXY species (X, Y = Cl, Br, OH). In clouds and marine aerosols, HgXY(aq) converts to HgCl₂ due to the abundance of Cl⁻ (Hedgecock and Pirrone, 2001).

We use a rate coefficient of $9 \times 10^{-14}$ cm³ mole⁻¹ s⁻¹ (Pal and Ariya, 2004) in a sensitivity simulation (see Section 4).

Table 2

<table>
<thead>
<tr>
<th>Chemical reactions and equilibria in the MBL model.</th>
</tr>
</thead>
<tbody>
<tr>
<td>Equation or equilibrium</td>
</tr>
<tr>
<td>HgII + O₃ → HgO + O₂</td>
</tr>
<tr>
<td>HgII + Br → HgBr</td>
</tr>
<tr>
<td>HgBr + HgO → Hg + Br</td>
</tr>
<tr>
<td>HgBr + Br → HgBr₂</td>
</tr>
<tr>
<td>HgBr + OH → HgBrOH</td>
</tr>
<tr>
<td>HgBr + Br → Hg + Br₂</td>
</tr>
<tr>
<td>HgCl₂ + Cl⁻ → HgCl⁻ + Cl⁻</td>
</tr>
<tr>
<td>HgCl₂(aq) + Cl⁻ → HgClBr(aq)</td>
</tr>
</tbody>
</table>

Species are gas-phase except where (aq) indicates aqueous-phase.

Equilibrium coefficients are in units of M atm⁻¹ or M⁻¹ as indicated. T is temperature in K. [M] is the number density of air in mol cm⁻³.

The mechanism and products of this reaction are poorly understood, and the reaction might be insignificant under atmospheric conditions (Calvert and Lindberg, 2005; Hayes et al., 2008). In our standard simulation, we assume a gas-phase HgO product with the same solubility as HgCl₂. Suppression of this reaction in the model can be accommodated by increasing the entrainment flux from the free troposphere as discussed in Section 4.

Assumed to be limited by the rate of HgCl formation, with Y representing a radical oxidant (Y = Cl, Br, OH) that subsequently adds to HgCl to form stable HgCl₂. The dissociation rate coefficient of HgCl, $k_1$, can be estimated from $\Delta G = -RT \ln(k_1/k_{1/2})$, where $\Delta G = -80.3 \text{ kJ mol}^{-1}$ is the Gibbs free energy of HgCl (Linstrom and Mallard, 2005).

RGM concentrations in the range 10–20 pg m⁻³ in the lower free troposphere over the regions examined here. If all of that HgII were RGM our model would overestimate mean MBL RGM concentrations as well as underestimate the relative amplitudes of diurnal cycles. We find a better match in our box model by entraining free-tropospheric RGM concentrations in the range 10–20 pg m⁻³ as given in Table 1. These values are within the wide range of the few free-tropospheric observations (Swartendruber et al., 2006; Sillman et al., 2007), and could be accommodated in the GEOS-Chem model if half of HgII were in particular form.
4. Model RGM diurnal cycles and budget

Fig. 1 shows simulations of RGM for each observed region. The model reproduces major features of all RGM data sets: nighttime minimum, early morning rise, and midday peak before 13 LST. The peak-to-peak diurnal amplitudes in the model are 0.4–3.5 times the 24-h mean, as compared to 1–3 in the field data, indicating that the model RGM lifetime is consistent with observations. The consistency in 24-h mean concentrations reflects the choice of entrained RGM concentration (Section 3). At Okinawa, the simulated mean exceeds measurements by a factor of 2. Since less entrainment would imply decreasing RGM with altitude and the relative diurnal cycle is underestimated by a factor of 2, we believe that the RGM sinks are underestimated. Sea-salt aerosols generated in the surf zone may explain the discrepancy.

We find that the RGM observations are most consistent with 2–9 × 10^6 atoms Br cm^-3 (24-h mean) over the midlatitude and subtropical Pacific and at Okinawa (Table 1). Using photochemical steady-state equations and O3 measurements, the corresponding peak BrO concentrations are 0.6 ppt in the Pacific midlatitudes, 0.8 ppt in the Pacific subtropics, and 1.2 ppt at Okinawa. These are within the large range of BrO observations in the MBL and consistent with higher BrO in coastal zones (Leser et al., 2003).

Simulations of the Atlantic data reproduce the relative diurnal amplitude (3.5 in the model vs. 3 in observations) but the maximum is too early in the day. Matching the diurnal amplitude requires 4.3 × 10^6 atoms Br cm^-3 (24-h mean), corresponding to 5.2 ppt BrO at sunrise. While there are few observations exceeding 2 ppt BrO in the remote MBL (Leser et al., 2003), the strong winds of the Atlantic cruise should have generated abundant sea-salt aerosols that might support such high Br levels. The strong winds also challenge the steady-state assumption used in the model.

Sea-salt aerosols contain 80–90% of the total MBL Hg0 in the model. That is 25–45 pg Hg0 m^-3 under most conditions. The simulation of the Atlantic data has up to 100 pg Hg0 m^-3 or 95% of Hg0 in aerosols. These values compare well with measurements of 5–40 pg Hg0 m^-3 (mean 20 pg m^-3) in coarse marine aerosol in Florida (Malcolm et al., 2003). Measurements in Alabama found somewhat lower values, 13 pg Hg0 m^-3, in total aerosol samples from airmasses originating over the Gulf of Mexico (Engle et al., 2008). Jaife et al. (2005) reported only 3 pg m^-3 of particulate Hg from the Okinawa site but their instrument excluded coarse aerosols and hence most sea-salt aerosol mass.

Fig. 4 gives steady-state 24-h average budgets of RGM sources and sinks in the model MBL for the Pacific and Okinawa cases. We do not present the Atlantic case because the model does not fit these observations well. From Fig. 4 we see that Hg0 + Br is the largest source of RGM, supplying 35–60%. Subsidence from the free troposphere is the next largest budget term, accounting for 25–40% of RGM in the MBL. Hg0 + O3 provides most of the remaining chemical source (5–20%). Hg0 + Cl plays a minor role (3–7%), but it could be larger if Hg0 + Cl kinetics are faster than assumed or if Cl concentrations are greater; this would imply a somewhat smaller contribution from Hg0 + Br.

One the sink side, 80–90% of MBL RGM enters the ocean. Uptake into sea-salt aerosols and their subsequent deposition contributes 65–80% of the total RGM loss for the Pacific cruise and Okinawa, while direct dry deposition of RGM contributes 10–15%. For the higher wind speeds that were observed in the Atlantic subtropics, the fraction of RGM lost through sea-salt aerosols rises to 90%. The RGM lifetimes range from 50 min (Atlantic subtropics) to 8 h (Okinawa).

Ventilation to the free troposphere provides a sink for 10–20% of RGM in the MBL partly compensating for entrainment. An entrainment velocity of 0.5 cm s^-1 at the top of a 750 m thick boundary layer gives a 40 h residence time for air in the MBL. Combined with the ~100 d chemical lifetime of Hg0 under most conditions, this implies that no more than 3% of Hg0 emitted into the MBL is oxidized before escaping to the free troposphere above. A corollary is that most of the Hg0 formed within the MBL has passed through the free troposphere as Hg0. Mercury deposition to the ocean is thus mainly derived from the global mercury pool, rather than from marine emissions in a closed MBL cycle.

The total chemical production of RGM in the model is 15–28 pg m^-3 d^-1, excluding the Atlantic. This is less than 2% loss of Hg0 per day and below the 4–8 pg m^-3 h^-1 limit which would produce an observable diurnal cycle. Thus our model is consistent with the observed absence of diurnal Hg0 variation without invoking a daytime Hg0 source to compensate for the photochemical Hg0 sink.

As mentioned in the Introduction, there are large uncertainties regarding gas-phase Hg chemistry, with most previous studies assuming OH and O3 to be the major Hg0 oxidants. Fig. 1 shows the diurnal cycle of RGM that would result from a fast Hg0 + OH reaction (Pal and Ariya, 2004) plus Hg0 + O3 and without Hg + halogen reactions. In this model RGM rises significantly only after 9 LST, due to the slow morning rise of OH (Nowak et al., 2001). In contrast, RGM simulated with halogen oxidants increases at sunrise, as it does in the observations. The RGM peak due to Hg0 + OH (14–15 LST) is delayed several hours relative to the observations and base simulation. Only when the apparent RGM lifetime is ~1 h or less, as in the subtropical Atlantic data, is the Hg0 + OH reaction consistent with the observed time of peak RGM. But reproducing the observed diurnal amplitude in the Atlantic data with the Hg0 + OH reaction still requires large additional photochemical RGM sources. Due to the deficiencies in the Atlantic simulation with oxidation by either OH or Br (described above), the Atlantic data alone do not resolve the major oxidant. Nevertheless, the majority of available RGM data are more consistent with Hg0 oxidation by Br, due to the earlier onset of bromine photochemistry. These tests cannot entirely reject the Hg0 + OH reaction from a model that also includes Hg + halogen reactions, but any significant contribution from OH would cause simulated RGM concentrations to lag behind the data.
While we have included the gas-phase Hg$_0^{a}$ + O$_3$ reaction (Hall, 1995) in our simulations, several investigators have challenged its mechanism or atmospheric relevance (Shepler and Peterson, 2003; Calvert and Lindberg, 2005; Hynes et al., 2008). This reaction is a small source of MBL RGM in our model (5–20%), and it has little diurnal variability, so its contribution could be replaced with a larger entrainment flux from the free troposphere. Such an increase would in fact better fit the observed and model values of RGM concentrations in the free troposphere.

5. Conclusions

We have analyzed atmospheric observations of reactive gaseous mercury (RGM) from Pacific and Atlantic cruises and from Okinawa in terms of mercury chemistry and deposition in the marine atmosphere. The dominant feature of the observations is a consistent diurnal cycle of concentrations with midday maximum, nighttime minimum, sharp increase after sunrise, and sharp decrease in afternoon. The early morning rise and midday peak implicate an RGM source from Hg$_0^{a}$ oxidation by halogen atoms. The relative diurnal amplitude implies an RGM lifetime of only a few hours. Nighttime RGM is anticorrelated with wind speed, pointing to a fast deposition sink. In contrast to RGM, Hg$_0$ shows no significant diurnal cycle in the observations.

We developed a box model for the marine boundary layer (MBL) to examine the consistency between the RGM data and current understanding of RGM sources and sinks, assuming that all gas-phase Hg$_0^{a}$ species are measured as RGM. After choosing concentrations of Br and free-tropospheric RGM within observational constraints to best match the amplitude and mean of the RGM diurnal cycle, we find that the observed cycle is consistent with a major RGM source from oxidation of Hg$_0^{a}$ by Br atoms. The implied MBL BrO concentrations are in the range 0.6–1.2 ppt, consistent with available data, although simulation of the Atlantic cruise data requires 5 ppt BrO. Oxidation of Hg$_0^{a}$ by Cl atoms and ozone is much smaller. A sensitivity simulation with Hg$_0^{a}$ oxidation mainly by OH generates a later rise and peak of RGM than observations, suggesting that the reaction is unimportant. The dominance of Br as a Hg$_0^{a}$ oxidant in the MBL was previously reported in box model calculations by Hedgecock et al. (2004, 2005), using older kinetic data for the Hg$_0^{a}$ + Br reaction (Ariya et al., 2002) that implied much faster oxidation than the more recent data used here (Goodsite et al., 2004). This resulted in daytime Hg$_0^{a}$ depletion inconsistent with observations unless a compensating diurnal cycle in oceanic emission was assumed. The slower rates used here yield a Hg$_0^{a}$ lifetime ~ 100 d in the MBL, consistent with the observed lack of Hg$_0^{a}$ diurnal variation.

Loss of RGM by dry deposition is too slow to account for the rapid decrease of RGM observed in the afternoon and evening hours. We find instead that the major loss pathway for RGM is uptake into sea-salt aerosols followed by aerosol deposition. We calculate the uptake rate with a physical mechanism accounting for sea-spray emission, aerosol deposition, aerosol hygroscopic growth and associated changes in Cl$^-$ concentration, and mass-transfer limitation. High Cl$^-$ concentrations present in sea-salt aerosols favor HgCl$_3$ and HgCl$_4^{2-}$ complexes and enhance Hg$_0^{a}$ solubility. Our model thus predicts that ~80% of MBL Hg$_0^{a}$ is present as sea-salt aerosol rather than RGM. The associated sea-salt aerosol concentrations of Hg$_0^{a}$ are 25–45 pg m$^{-3}$, consistent with the few observations available. Uptake by sea-salt aerosols accounts for 65–80% of RGM loss from the MBL in our simulation, and enables us to generally reproduce the observed phase and diurnal amplitude of the RGM observations. Considering direct dry deposition as well, 80–90% of RGM in the MBL deposits to the ocean.

We calculate a ~100 d lifetime for Hg$_0^{a}$ in the MBL against oxidation to Hg$_0^{a}$ in most conditions. Since ventilation of the MBL to the free troposphere takes place on a time scale ~ 1 d, the long Hg$_0^{a}$ lifetime implies that Hg$_0^{a}$ emitted from the ocean is ventilated to the free troposphere rather than oxidized in the MBL and returned to the ocean in a closed system. A corollary is that RGM formed and deposited within the MBL derives primarily from Hg$_0^{a}$ transported in the free troposphere over long distances in the global mercury cycle.

Entrainment of RGM from the free troposphere in our model accounts for 25–40% of RGM in the MBL, second in importance to Hg$_0^{a}$ oxidation by Br atoms. Selin et al. (2008) previously proposed that subtropical subsidence of high-Hg$_0^{a}$ air would be a major contributor to oceanic deposition of mercury. It provides a steady source of RGM to the MBL that our model requires to account for the observed magnitude of MBL concentrations and their relative diurnal amplitude. It is also consistent with the few RGM observations in the lower free troposphere.

Several key parameters and processes in our understanding of MBL mercury cycling have large uncertainties that require further study. Perhaps most important are measurements of BrO in the MBL with sub-ppt sensitivity, to test the hypothesis that Br is the dominant Hg$_0^{a}$ oxidant. Mercury observations under intermediate winds (6–10 m s$^{-1}$) are also necessary in light of our poor understanding of the RGM diurnal profile shape under strong winds. Vertical profiles of Hg$_0^{a}$ from the sea surface to the free troposphere, especially in the subtropical subsidence zones, would better constrain entrainment. Finally, Hg$_{aq}^{+}$ concentrations in sea-salt aerosols should be measured to test the hypothesis that sea-salt aerosol deposition is the dominant pathway for mercury deposition to the ocean.

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